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Dedicated to the memory of Alexei Starobinsky

Abstracts

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18-й Российская гравитационная конференция – Международная конференция по гравитации, космологии и астрофизике «RusGrav-18». Тезисы докладов. — Казань: Изд-во КФУ, 2024. – 122 с.

Сборник содержит тезисы докладов участников 18-й Российской гравитационной конференции – Международной конференции по гравитации, космологии и астрофизике «RusGrav-18». В материалах представлены работы ведущих специалистов и начинающих исследователей из научных центров России, ближнего и дальнего зарубежья, посвящённые исследованиям в ОТО и модифицированных теориях гравитации, теоретической и наблюдательной космологии, релятивистской астрофизике. Международная конференция проходила в Казанском федеральном университете с 25 по 29 ноября 2024 года.

Сборник адресован научным работникам, аспирантам и молодым ученым, специализирующимся в области теории гравитации, космологии и астрофизики, а также для студентов старших курсов естественнонаучных направлений.

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BLACK HOLE UNIQUENESS THEOREMS IN GAUGED SUPERGRAVITIES

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The classification of anti de Sitter black holes is an open problem of central importance in holography. In this talk, I will present new advances in classification of supersymmetric solutions to five-dimensional gauged supergravity coupled to arbitrary number of Abelian vector multiplets. In particular, we prove a black hole uniqueness theorem within a subclass of solutions with biaxial symmetry and a certain twistor structure. This subclass includes all known solutions of this theory, both analytic and numerical.

The logic of the proof is based on the near-horizon data only and is agnostic of the asymptotic region. We, therefore, deduce that a classification of CFT states in the UV of the holographic correspondence is possible to derive by deep IR arguments.

BLACK HOLES IN SEMICLASSICAL $f(R)$ GRAVITY WITH EXTRA DIMENSIONS

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In this work we obtained the solution of the equations of the semiclassical $f(R)$ theory of gravitation ($c = \hbar = 1$)

$$-\frac{1}{2}f(R)\delta_A^B + (R_A^B + \nabla_A \nabla^B - \delta_A^B \square) \frac{df(R)}{dR} = \frac{8\pi}{m_D^2} \langle T_A^B \rangle, \quad (1)$$

describing a black hole in $(4 + n)$ -dimensional spacetime

$$ds^2 = \left(1 - \frac{2M}{r}\right) dt^2 - \frac{dr^2}{\left(1 - \frac{2M}{r}\right)} - r^2 (d\theta^2 + \sin^2 \theta d\varphi^2) - L_0^2 d\Omega_n^2, \quad (2)$$

where M is the integration constant, $A, B, C, \dots = 1, 2, \dots, 4 + n$, $f(R)$ is an arbitrary function of scalar curvature R , $m_D - (4 + n)$ - dimensional Planck mass, ∇_A - denotation of the covariant derivative, R_A^B - Ricci tensor, $\square = g^{AB} \nabla_A \nabla_B$. The additional space is assumed to be an n - dimensional sphere of radius L_0 , much smaller than the characteristic scale of curvature l ($L_0 \ll l$) of the four-dimensional part of space

The right part of equations (1) describes the expectation value of the stress-energy tensor operator of the quantized matter fields. In the limit $l_{\text{Pl}} \ll L_0$ the structure of this tensor has the form (see Popov A., Phys. Rev. D (2001), V. 64, 104005)

$$\langle T_A^B \rangle = \frac{l_{\text{Pl}}^2}{L_0^4} K_A^B \left(1 + O(L_0^2/l(r)^2) \right), \quad (3)$$

where in the coordinates (2)

$$K_A^B = \text{diag} \left(K_t^t, K_t^t, K_t^t, K_t^t, K_5^5, \dots, K_5^5 \right), \quad (4)$$

and K_t^t and K_5^5 are constant values. The solution (2) is obtained in the case of neglecting the small terms of the expansion (3) and is possible with the following relations between parameters $R(r) = R_0, L_0, f(R_0)$ and $\left. \frac{df}{dR} \right|_{R=R_0}, \left. \frac{df}{dR} \right|_{R=R_0}$

$$f(R_0) = -\frac{16\pi}{m_D^2} \frac{R_0^2 K_t^t}{n^2(n-1)^2}, \quad (5)$$

$$\left. \frac{df}{dR} \right|_{R=R_0} = \frac{8\pi}{m_D^2} \frac{R_0 (K_5^5 - K_t^t)}{n(n-1)^2}, \quad (6)$$

$$L_0^2 = \frac{n(n-1)}{R_0}. \quad (7)$$

In the case of quadratic gravity $f(R) = aR^2 + R + c$ these expressions will take the form

$$L_0^2 = \frac{n(n-1)}{R_0} = \frac{8\pi}{m_D^2} \frac{(K_5^5 - K_t^t)}{(n-1)} - 4a, \quad (8)$$

$$c = \frac{(n-1)^2 \left(an^2(n-1)^2 + \frac{8\pi}{m_D^2} (K_t^t(n-2) - nK_5^5) \right)}{4 \left(an(n-1)^2 + \frac{4\pi}{m_D^2} (K_t^t - K_5^5) \right)^2}. \quad (9)$$

Note that taking into account the following terms of the expansion (3) leads to the dependence of the radius of the extra space L_0 , as well as the four-dimensional Planck mass and the four-dimensional cosmological constant (under the adopted assumptions, the theory can be reduced to Einstein's four-dimensional theory of gravitation) on the radial coordinate.

SPONTANEOUS BRANES FORMATION

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