

$\langle\varphi^2\rangle$ of a quantized scalar field in a nonzero temperature thermal state at a long throat

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An analytical approximation of $\langle\varphi^2\rangle$ for a quantized scalar field in a thermal state at arbitrary temperature is obtained. The scalar field is assumed to be both massive and massless, with an arbitrary coupling ξ to the scalar curvature, and in a thermal state at (an) arbitrary temperature T . The gravitational background is assumed to be static spherically symmetric

$$ds^2 = -f(\rho)dt^2 + d\rho^2 + r^2(\rho)(d\theta^2 + \sin^2\theta d\varphi^2). \quad (1)$$

and slowly varying

$$\varepsilon_{\text{WKB}} = L_\star/L \ll 1, \quad (2)$$

where

$$L_\star(\rho) = \frac{r(\rho)}{\sqrt{2\xi + m^2 r^2(\rho)}}, \quad (3)$$

and L is a characteristic scale of variation of $f(\rho)$ and $r^2(\rho)$:

$$\frac{1}{L(\rho)} = \max \left\{ \left| \frac{r'}{r} \right|, \left| \frac{f'}{f} \right|, \left| \frac{r'}{r} \sqrt{|\xi|} \right|, \left| \frac{f'}{f} \sqrt{|\xi|} \right|, \left| \frac{r''}{r} \right|^{1/2}, \left| \frac{f''}{f} \right|^{1/2}, \dots \right\}. \quad (4)$$

The result of the calculation is (A. Popov , Phys. Rev. D **94**, 124033 (2016))

$$\begin{aligned} 4\pi^2 \langle\varphi^2\rangle_{\text{ren}} &= \\ \frac{4\mu^2}{r^2} J(a/\mu) - \frac{1}{4r^2} \left(2\xi - \frac{1}{4}\right) + \frac{1}{4r^2} \left[m^2 r^2 + 2\left(\xi - \frac{1}{6}\right)\right] \ln \left| \frac{m^2 r^2 + 2\xi - 1/4}{m_{\text{DS}}^2 r^2} \right| \\ - \frac{\mu^2}{r^2} I_1(\mu) + \frac{r^{2''}}{24r^2} - \frac{r^{2'^2}}{48r^4} + \frac{f'r^{2'}}{48fr^2} + \left[\frac{f''}{12f} + \frac{r^{2''}}{6r^2} - \frac{f'^2}{24f^2} - \frac{r^{2'^2}}{24r^4} + \frac{f'r^{2'}}{12fr^2} \right. \\ \left. + \xi \left(-\frac{f''}{2f} - \frac{r^{2''}}{r^2} + \frac{f'^2}{4f^2} + \frac{r^{2'^2}}{4r^4} - \frac{f'r^{2'}}{2fr^2} \right) \right] \left(\frac{1}{2} \ln \left| \frac{m^2 r^2 + 2\xi - 1/4}{m_{\text{DS}}^2 r^2} \right| - \mu \frac{dI_0(\mu)}{d\mu} \right), \end{aligned} \quad (5)$$

where

$$\begin{aligned} a &= \frac{2\pi Tr}{\sqrt{f}}, \quad \mu^2 = m^2 r^2 + 2\xi - \frac{1}{4}. \\ J(a/\mu) &= \int_1^\infty \frac{\sqrt{\eta^2 - 1}}{e^{2\pi\mu\eta/a} - 1} d\eta, \quad I_m(\mu) = \int_0^\infty \frac{\eta^{2m-1} \ln|1 - \eta^2|}{1 + e^{2\pi\mu\eta}} d\eta. \end{aligned} \quad (6)$$

The constant m_{DS} is equal to the mass m of the field for a massive scalar field. For a massless field m_{DS} is an arbitrary parameter due to the infrared cutoff in $\langle\varphi^2\rangle_{\text{DS}}$. A particular choice of the value of m_{DS} corresponds to a finite renormalization of the coefficients of terms in the gravitational Lagrangian and must be fixed by experiment or observation.

We have shown that in a long throat the effect of vacuum polarization of a quantized scalar field is local one for any finite mass m of the quantized field, including $m = 0$.

The work is performed according to the Russian Government Program of Competitive Growth of Kazan Federal University.

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